

TFOP SG1

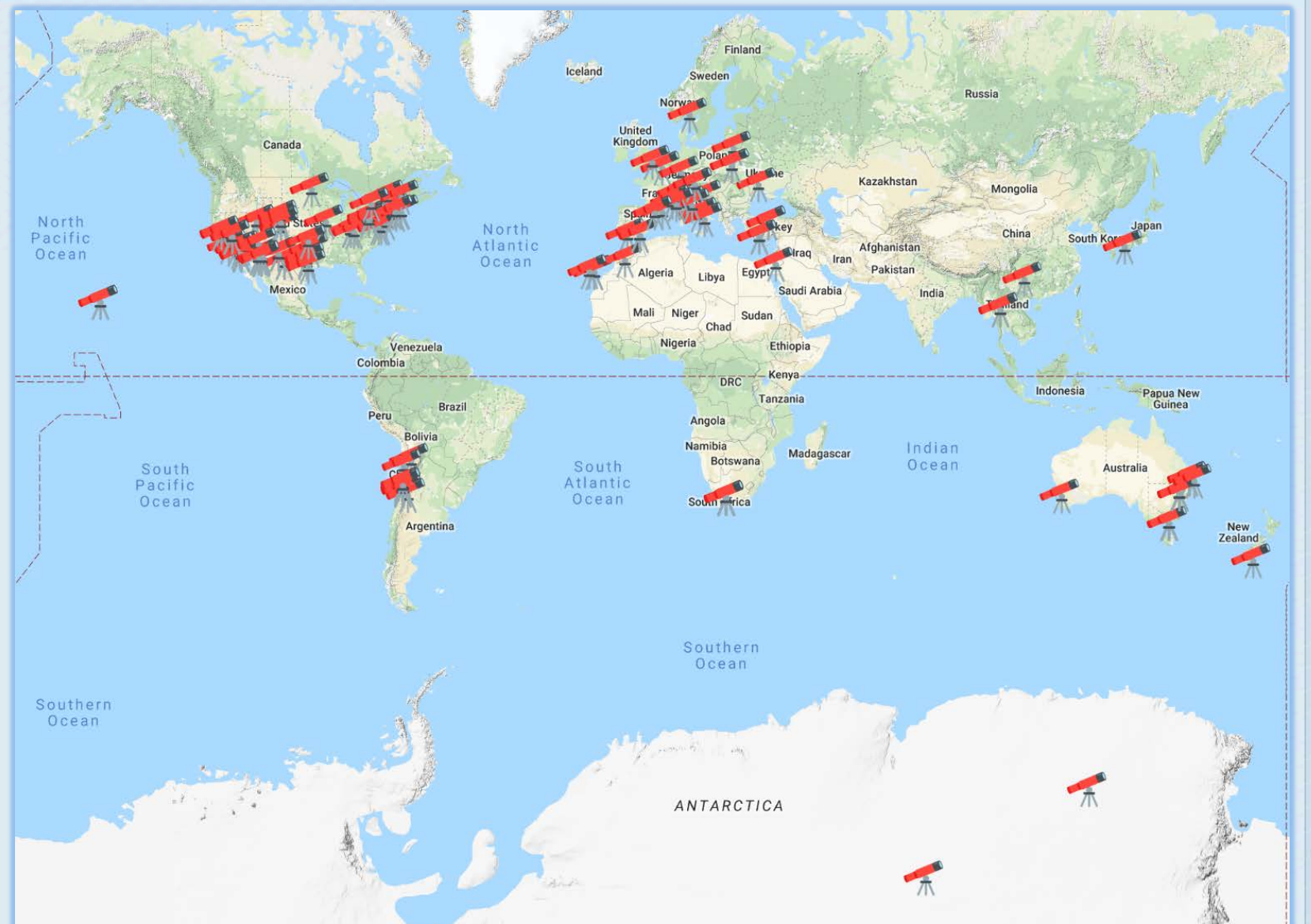
Ground-based Time-series Photometry: Goals, Status, Results and Your TESS Paper

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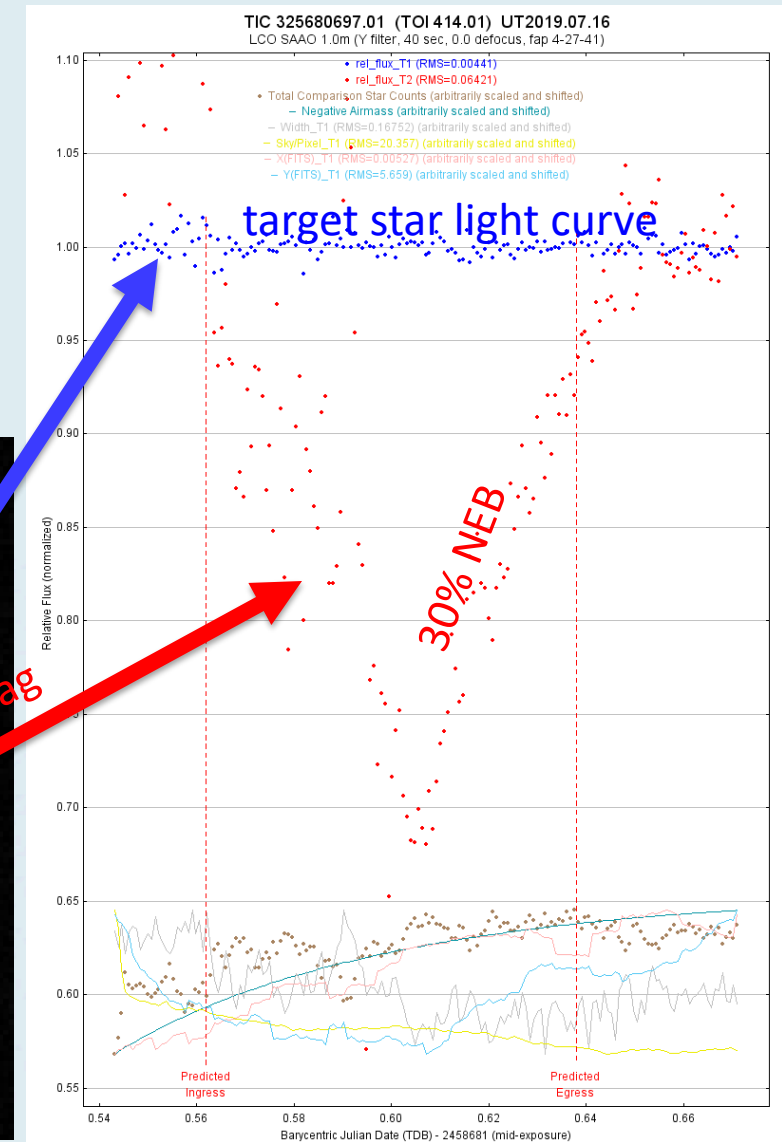
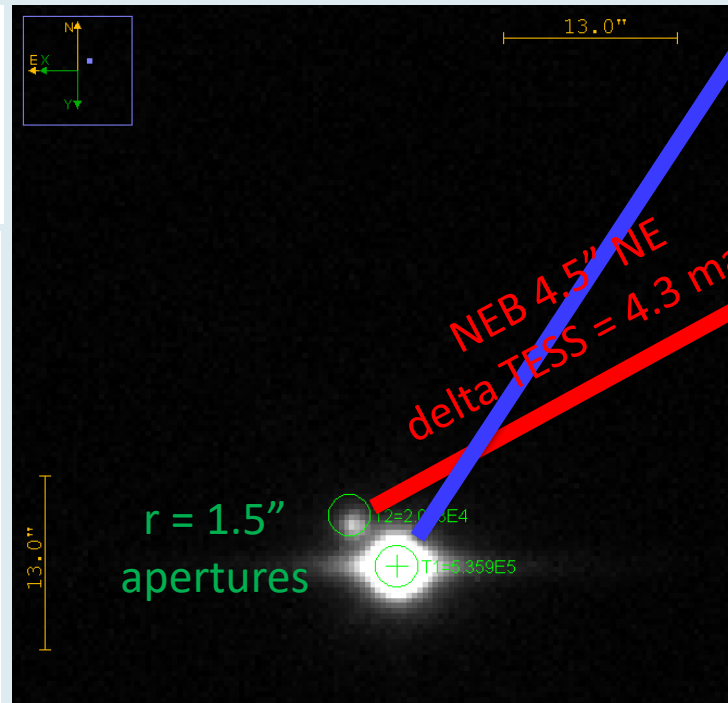
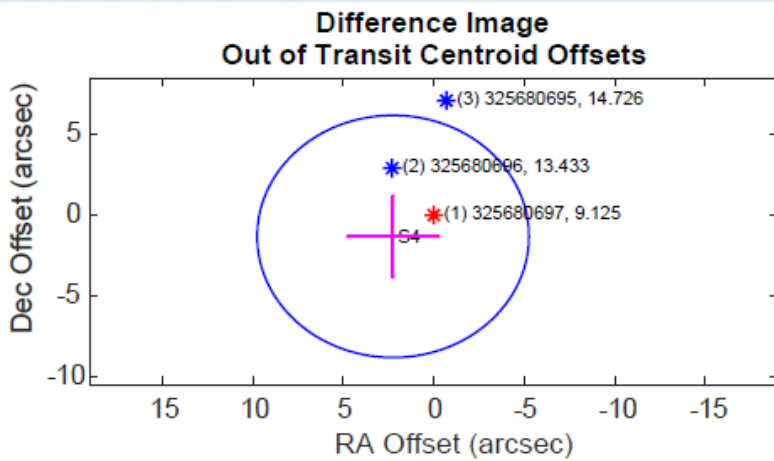
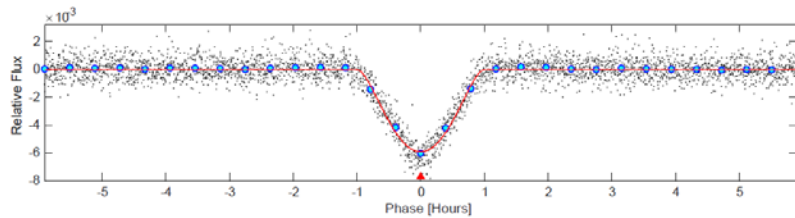
- ◆ SG1 Team is ~200 members strong
 - *Professional astronomers*
 - *Student astronomers*
 - *Citizen astronomers*
 - *Ground-based transit surveys*



◆ Goals:

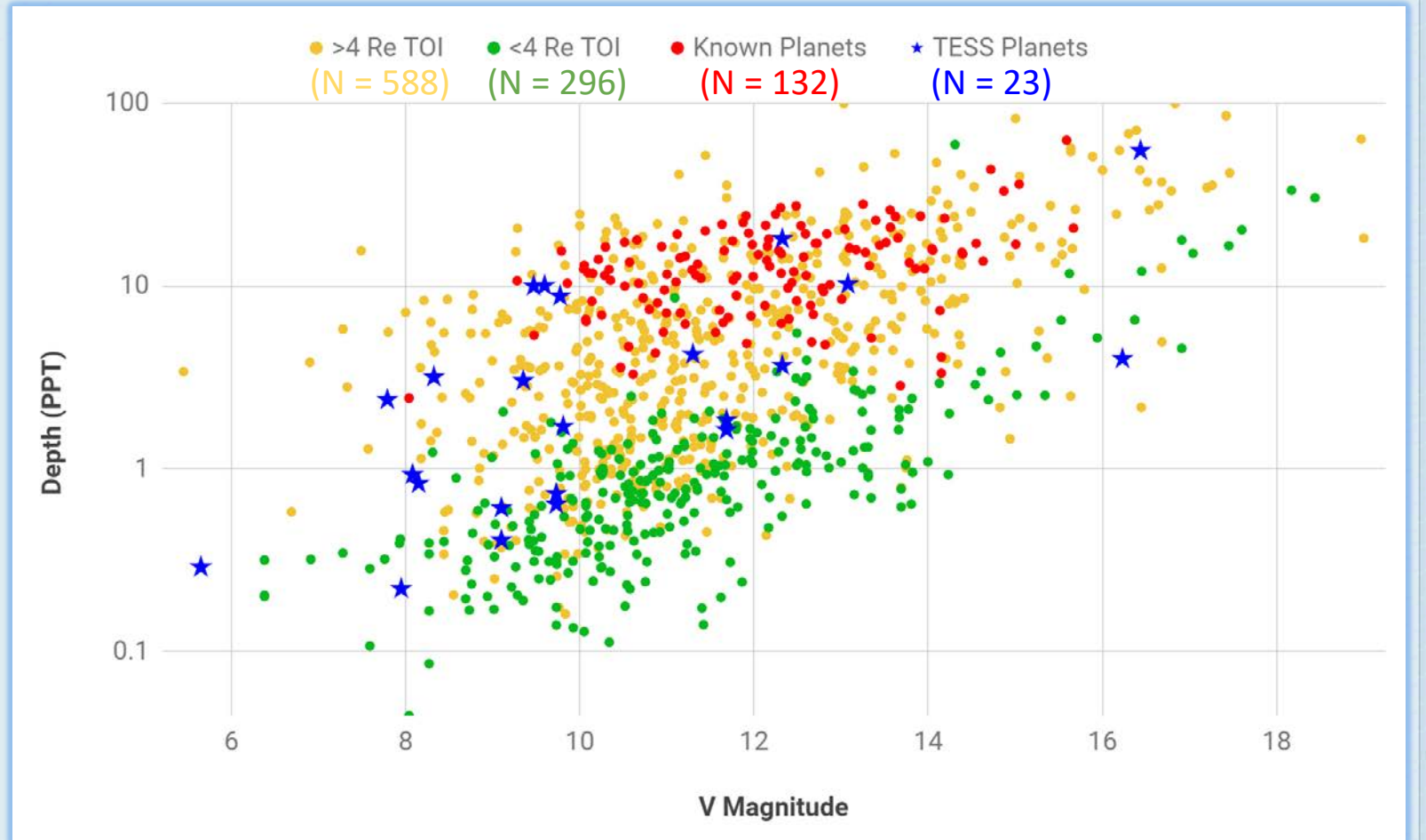
- *Identify photometric false positives*
- *Confirm events on-target*
- *Refine ephemerides*
- *Measure TTVs*
- *Measure transit depths to confirm TESS deblending factor*
- *Verify transit depth achromaticity*

- ◆ Search stars within 2.5 arcmin of target for nearby EBs (NEBs)
- ◆ If found 1st time => SG1 disposition **PNEB**
- ◆ If found 2nd time => SG1 disposition **NEB**



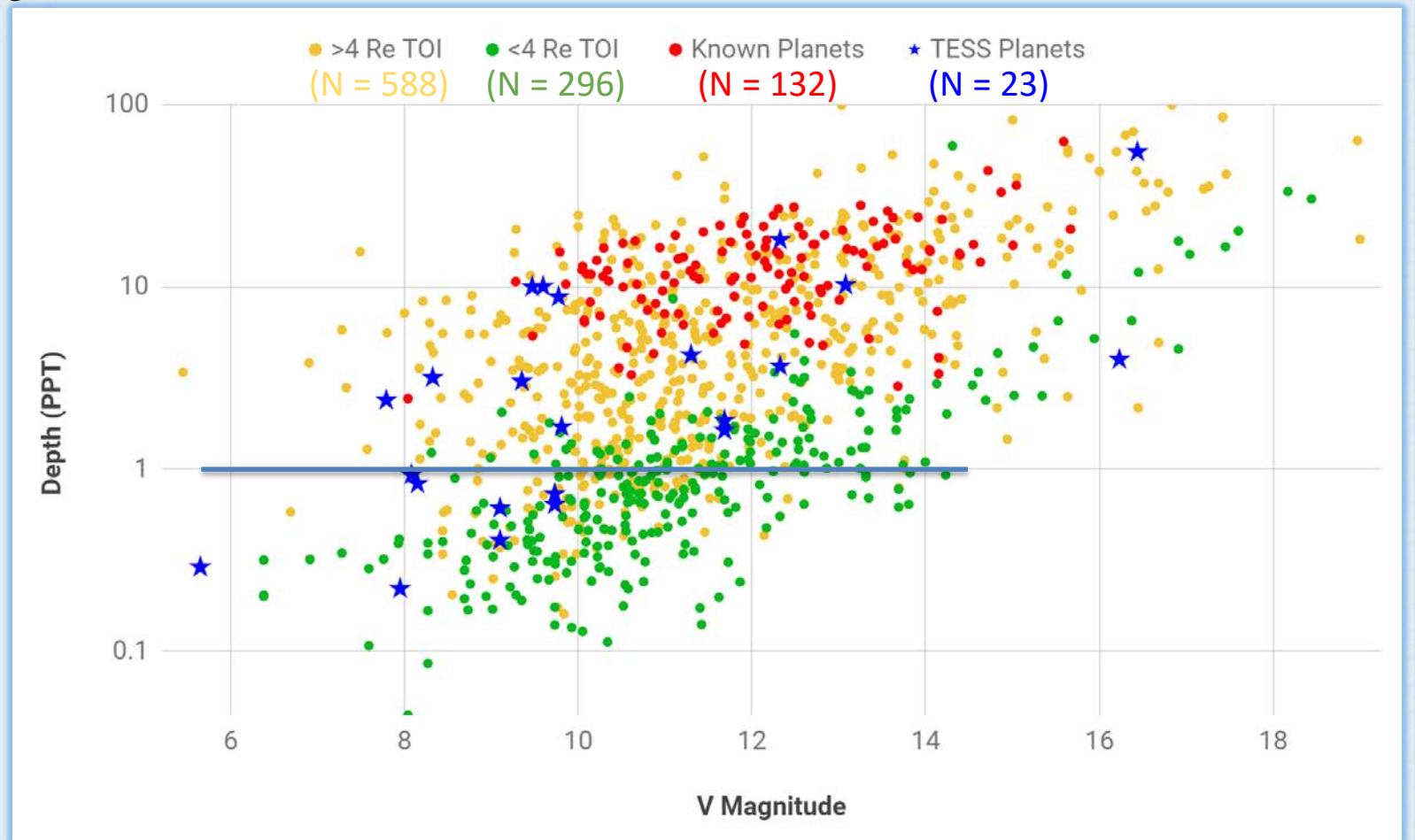
SG1 Goals are dependent on transit depth

- ◆ Ground-based sensitivity general threshold
 - ~ 1000 ppm (1 ppt or ~ 1 mmag)
- ◆ 40% of $<4 R_E$ TOIs > 1 ppt
- ◆ 75% all TOIs > 1 ppt
- ◆ Can also contribute at < 1 ppt



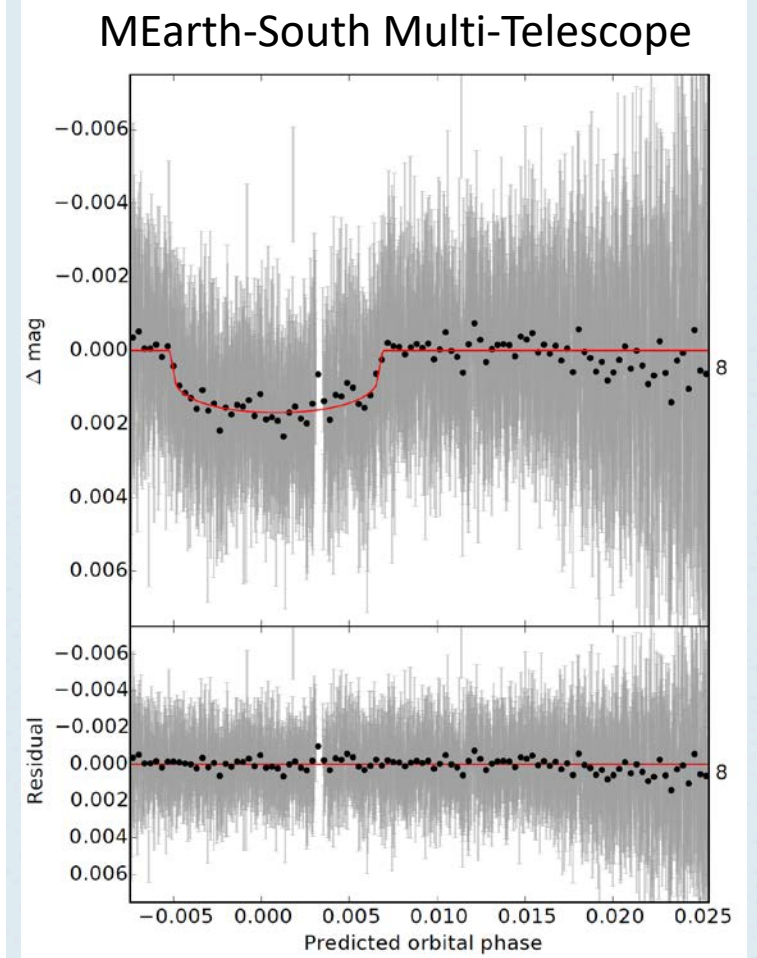
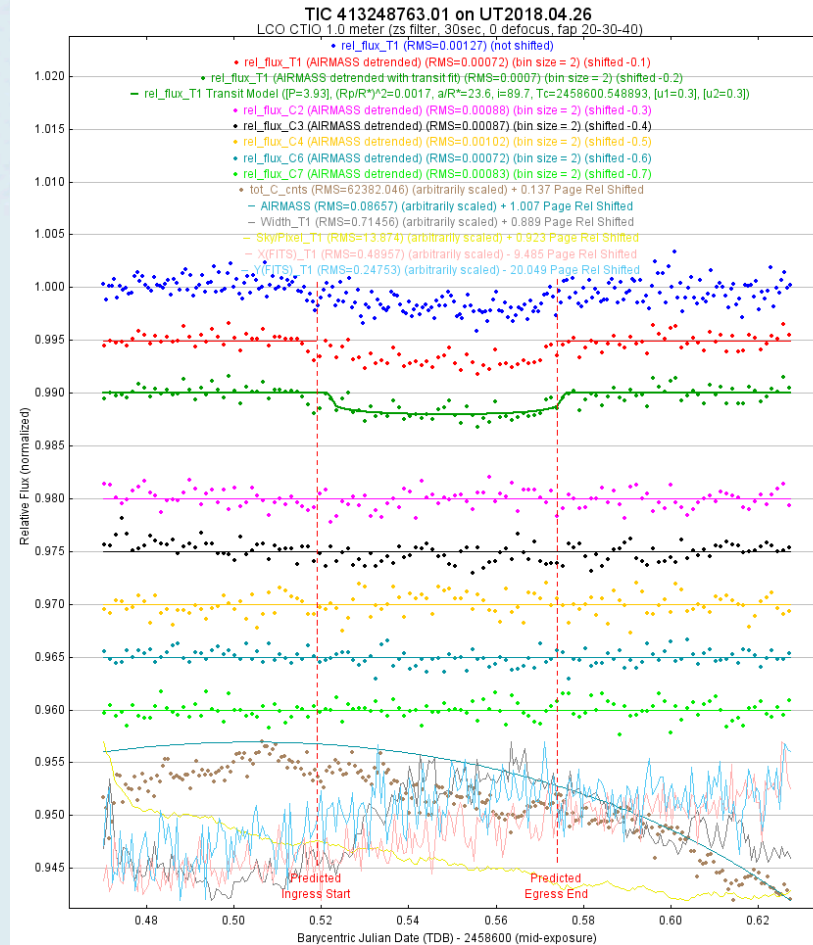
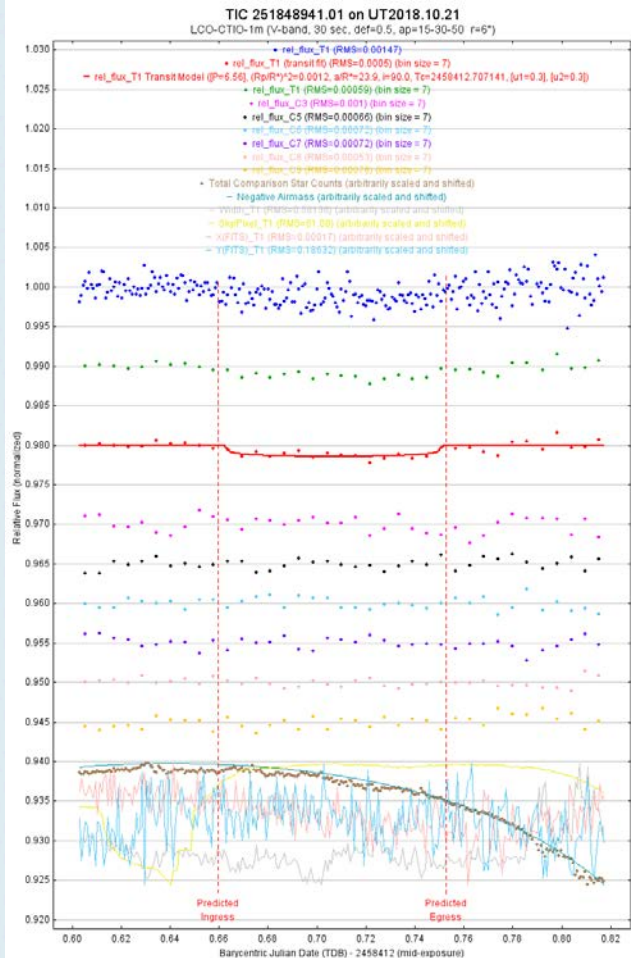
SG1 Goals are transit depth dependent

- ◆ Ground-based sensitivity general threshold
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- ◆ Can also contribute at < 1 ppt



- ◆ For transit depths > 1 ppt
 - *Primary goal is to confirm transit on target*
 - *Confirmed and no follow-up aperture contamination*
=> **VPC** = *Verified Planet Candidate (SG1 disposition)*
 - *Confirmed but follow-up aperture is contaminated with known nearby star*
=> **VPC-**
 - *Confirmed off target*
=> **NEB** = *nearby eclipsing binary*
 - *Secondary goal is a depth chromaticity check*
 - *No significant filter dependent depth*
=> **VPC+**
 - *Significant filter dependent depth*
=> **BEB** = *blended (in the follow-up aperture) EB false positive*

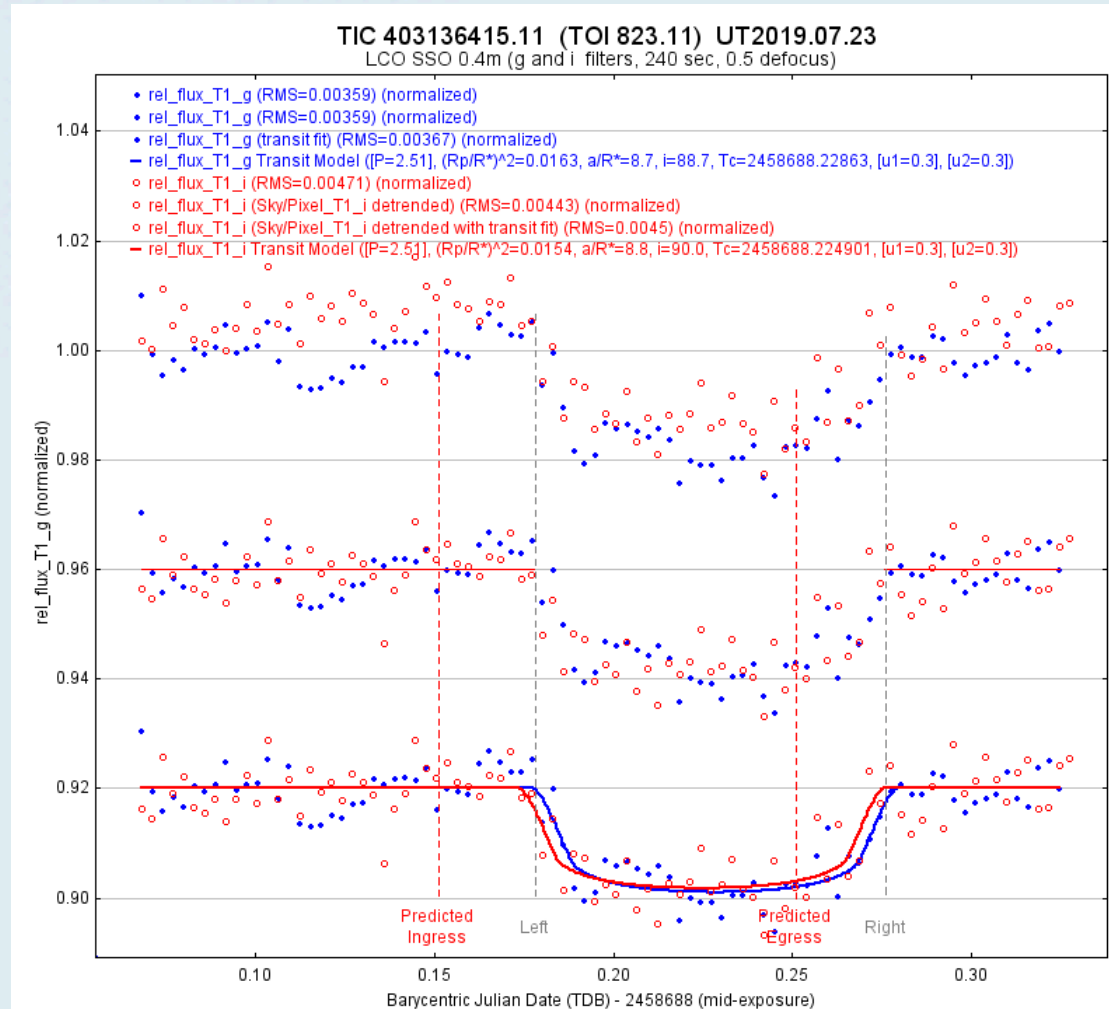
Transits detected on target => VPC disposition (often with 1-2" apertures)



Credit: Kevin Collins, Vanderbilt University/George Mason University

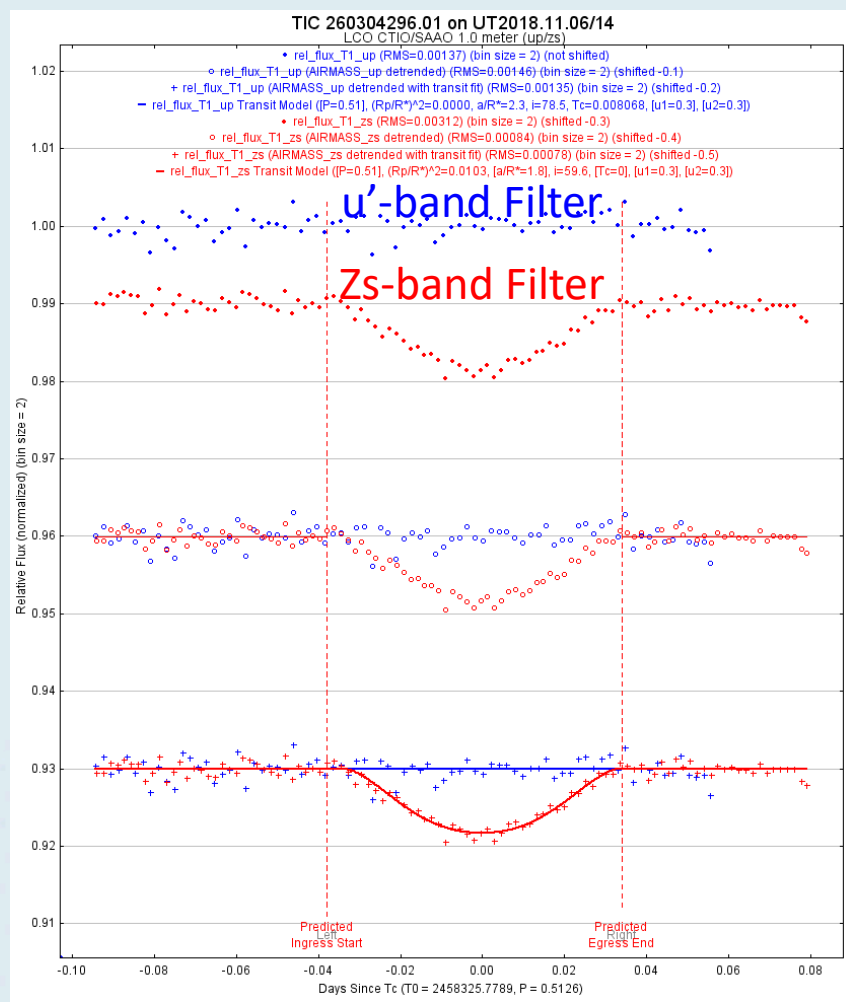
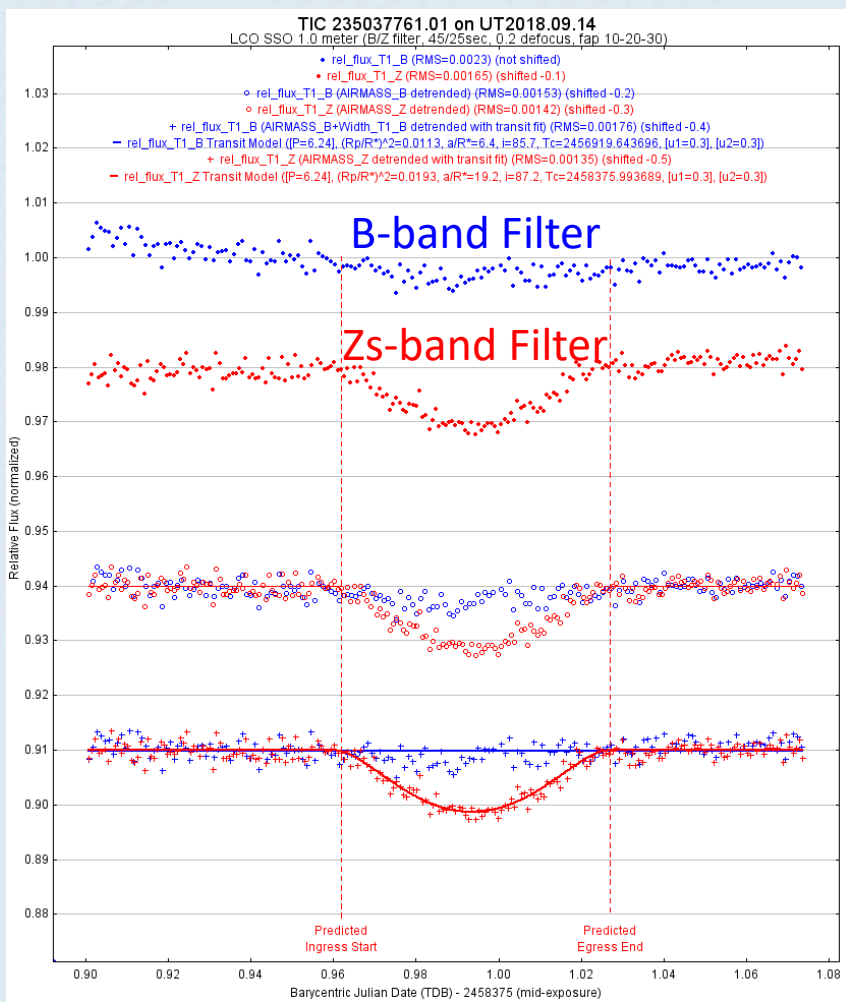
Credit: Jonathan Irwin

Achromatic Transits detected on target => VPC+ disposition



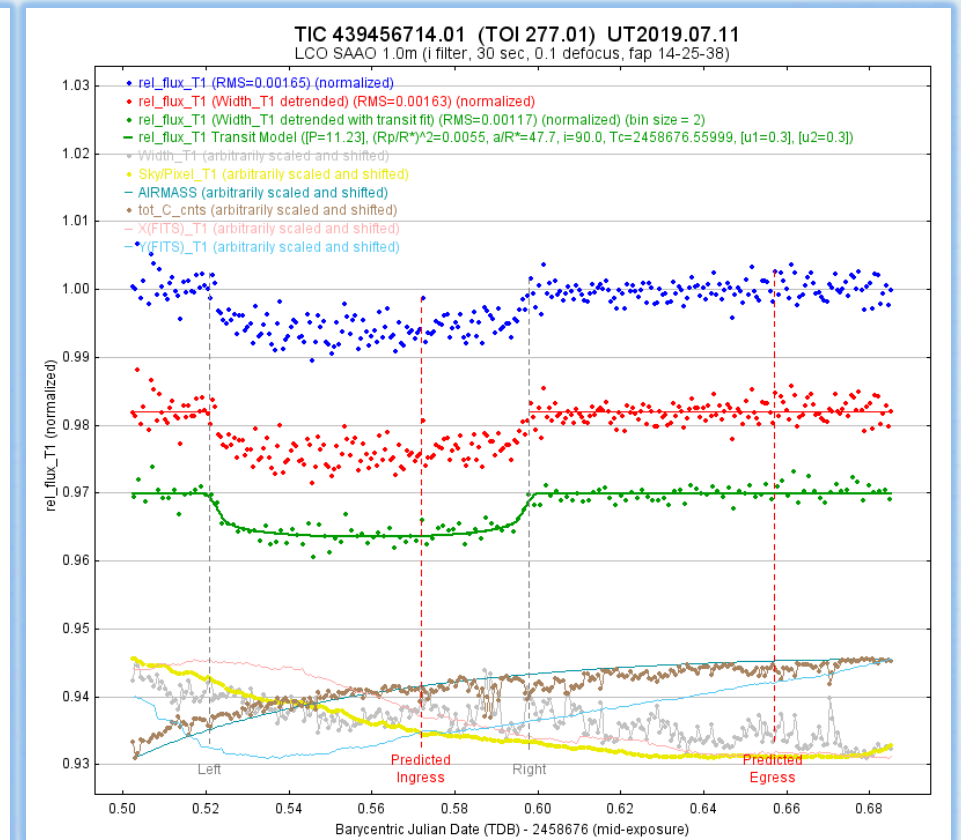
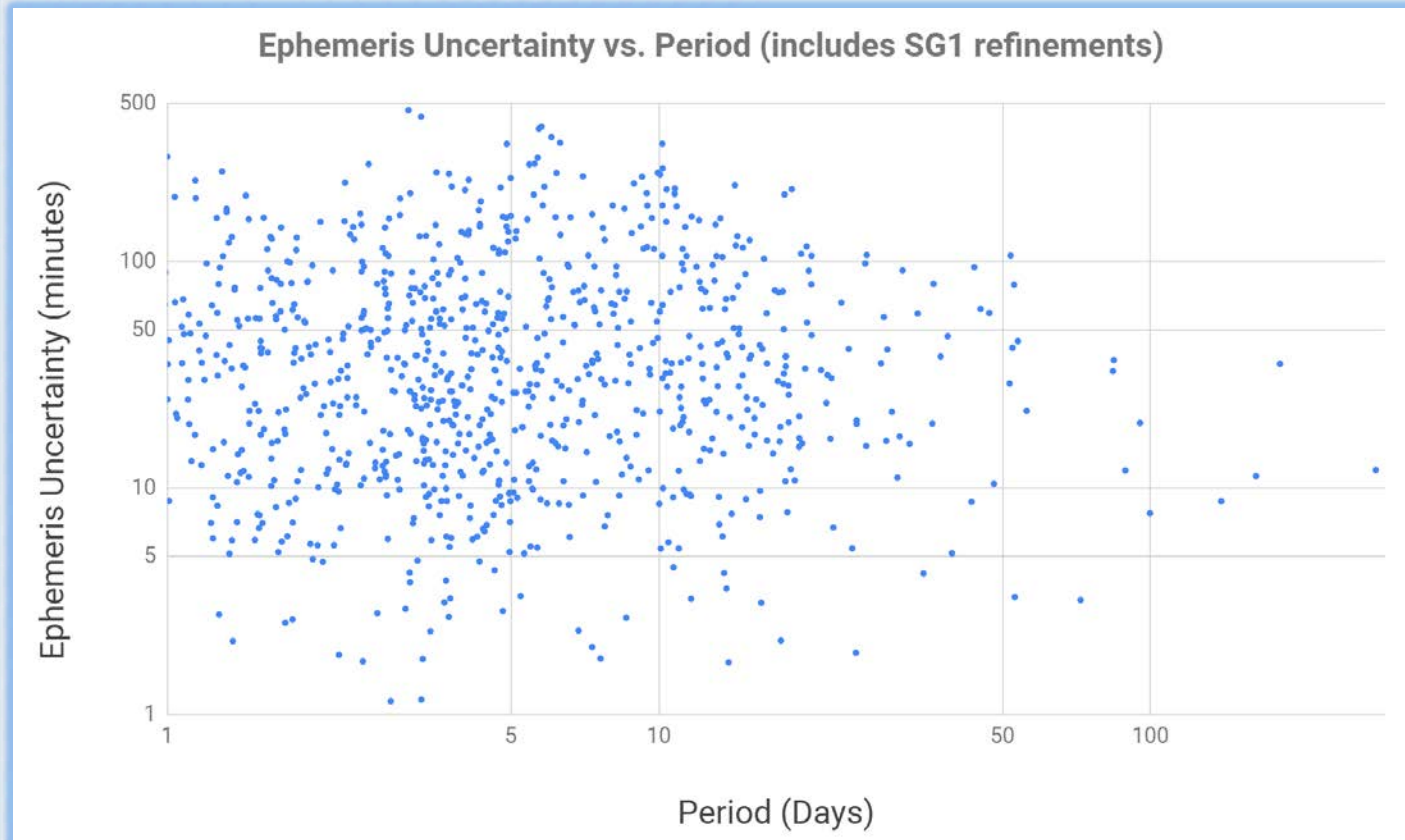
Credit: Dennis Conti, Citizen Astronomer

Examples of strong chromaticity detections => BEB False Positive

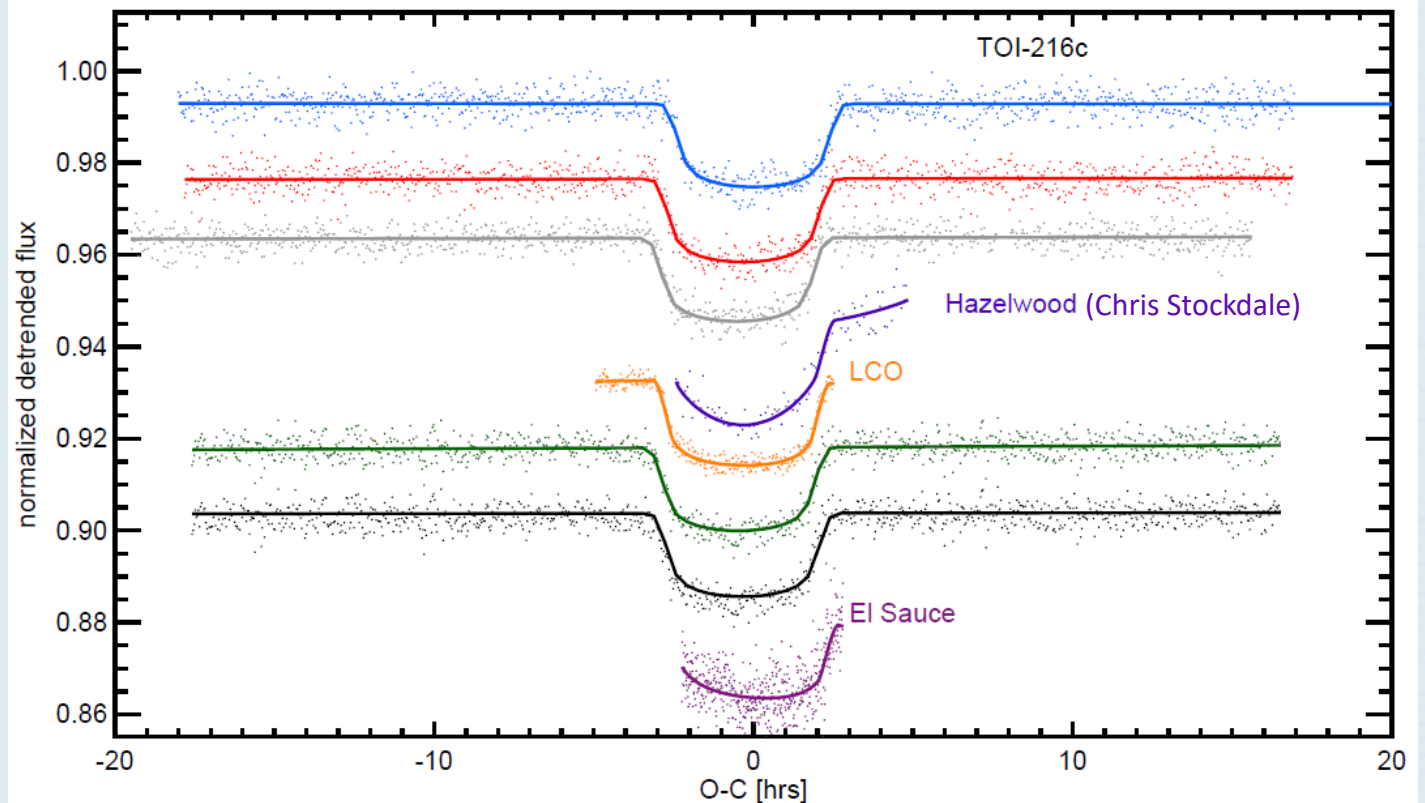
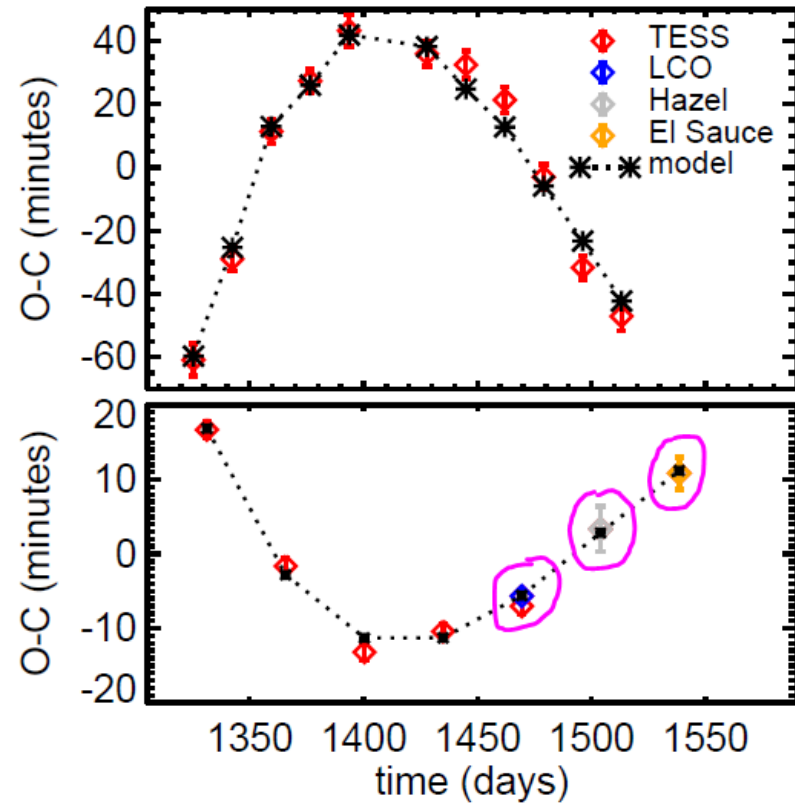


Credit: Kevin Collins, Vanderbilt University

- ◆ Several TOIs already have > 2-hour uncertainty
- ◆ Ephemeris is refined with each SG1 observation
- ◆ Uncertainty generally reduced by ~90%
- ◆ Chart below includes SG1 refinements



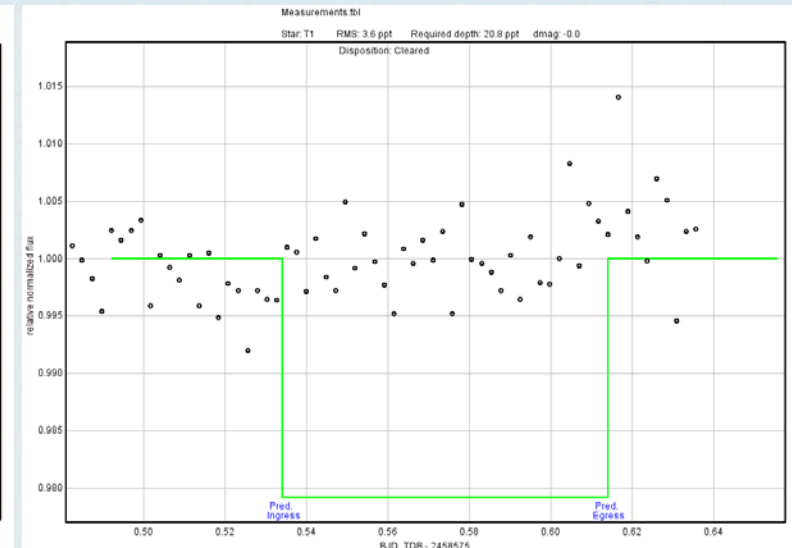
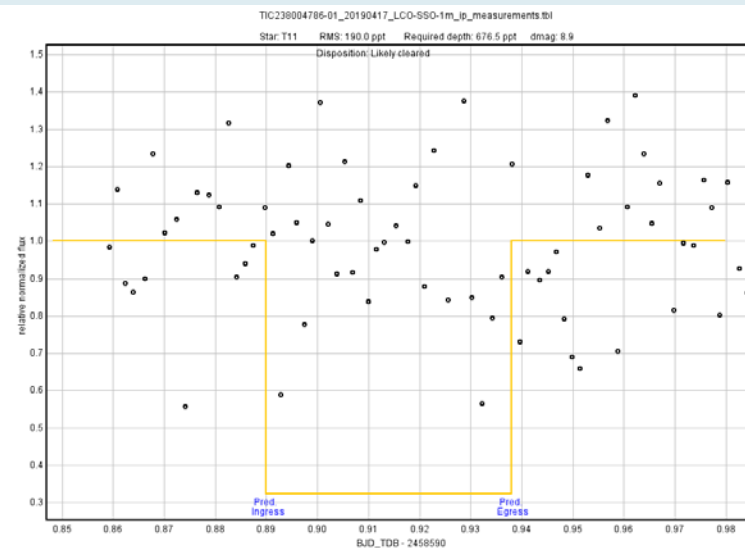
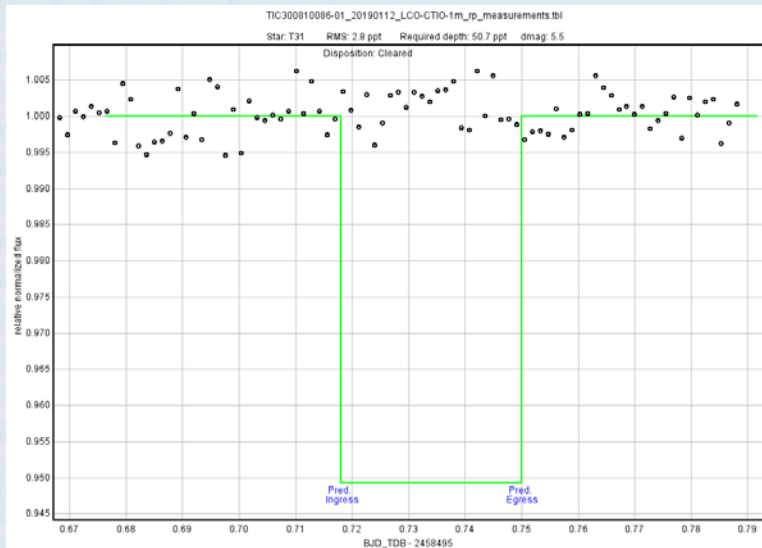
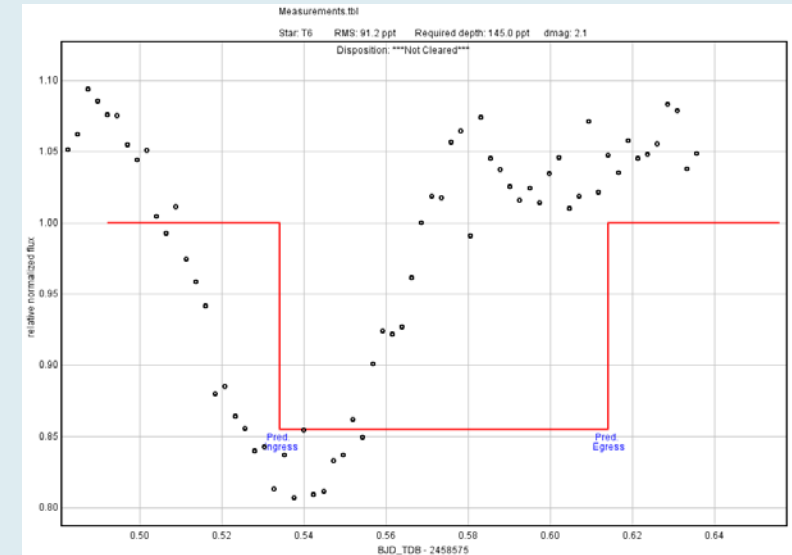
- ◆ TOI-216b and TOI-216c system
- ◆ SG1 citizen astronomers contributed to TTV measurements



Dawson et al. 2019

- ◆ For transit depths < 1 ppt
 - *Primary goal is to check all Gaia DR2 star locations within 2.5 arcmin of the target for NEBs*
 - *Confirmation of no NEBs (no Gaia stars blended in target PSF)*
*=> **CPC** = Cleared Planet Candidate*
 - *Confirmed but target star PSF is contaminated with known nearby star*
*=> **CPC-***
 - *Confirmed off target (1st time)*
*=> **PNEB** = possible nearby eclipsing binary*
 - *Confirmed off target (2nd time)*
*=> **NEB** = nearby eclipsing binary*

- ◆ Light curves of all stars within 2.5 arcmin of target are available on ExoFOP-TESS
- ◆ Box transit models show depth needed to produce TESS detection



- ◆ Any TFOP member can request access (send me your google affiliated email address)
- ◆ Contains all SG1 TOIs
 - *All Mission TOIs*
 - *Various special request candidates*
- ◆ Current SG1 priority
- ◆ Current SG1 disposition
- ◆ Observing history
- ◆ Current ephemeris uncertainty
- ◆ NOTE: all confirmed FPs are listed on ExoFOP-TESS

Target	TIC	TOI	Pipe-line	Sec-tor	Priority	Dis-po-sition	Vmag	Depth (ppm)	P	Durat-ion (hrs)	Rp (R_E arth)	Publication Status	N Obs	Comments	Cur Eph Unc (min)
25375553.01	25375553	143.01	SPOC	S1	5	KP	10.251	6939	2.3109	3.422	17.6		0	WASP-111	11
393940766.01	393940766	148.01	QLP	S1	3	VPC	12.398	5950	4.8661	3.215	9.7		4	Peter Nelson observed a full transit on 20180914 in V and found a tentative detection of the transit on target. Bob Massey observe a full transit on 20180914 in V and found a deeper transit, but with 14 ppt RMS. Fran and the TRAPPIST-South team observed a full transit on 20180928 in z' with an r=4.5" aperture. A ~5 ppt event was detected on target. No Gaia stars are contaminating the follow-up aperture. TG Tan observed an inconclusive ingress on 20181121 in V. The next observation should be a full transit in g/B to check for a chromatic depth.	92
260985861.01	260985861	149.01	QLP	S1	5	VPC	10.727	5082	3.3402	4.077			4	Giovanni Isopei and Franco Mallia observed an ingress on 20181001 and found an apparent on-time ~3 ppt ingress on target in an r=8" aperture. That depth should be considered a rough fit and possibly consistent with the TESS depth. No obvious potential NEBS were found down to ~ dmag~7.5. See Giovanni's dmag + RMS analysis for details. Howie Relles also observed this TOI and with in-transit and partial egress coverage. There is an apparent egress starting at about the right time and with a depth of at least 5 ppt. Fran and the TRAPPIST team observed a full transit on 20181114 in Rc and found a 5 ppt ~on-time event. Chris Sockdale observe in egress+50% and found an apparent ~10 ppt evt on target (but depth is uncertain). Gaia reports this system has two almost equal brightness neighbors that are separated by 1" and with dmag=0.1, so the depth would be ~twice what is reported by TESS. Gaia does not give a radius for either star. A SOAR speckle image on ExoFOP confirms the 1" neighbor. Retired as a VPC, but note that the target is two stars.	65
271893367.01	271893367	150.01	QLP	S1	5	VPC+	11.39	7386	5.8576	5.818	15.0	Kossakowski TOI-150b TOI-163b In Prep Canas non-TESS TOI-150.01 Accepted	3	Kevin Collins reduced an LCO-CTIO-1m egress + 60% on 20181109 in Zs and found a 6.6 ppt event on time and on target with an r=5.84" aperture which is not contaminated by any Gaia stars. Howie Relles observed the same partial event in ip from an LCO-CTIO-1m and found a ~6 ppt event on time and on target within a 19.5" aperture. Fran observed a full transit with meridian flip in u' on 20181219 and found a ~5 ppt event on target using a 4.6" aperture. The meridian flip and short post-egress baseline could make the model depth uncertain and alternate models give 6.5 and 7.5 ppt. Phil Evans observed a full in B on 20190130 in B and found a 6 ppt event on target in a 7.35" aperture. No more observations needed.	40
261136679.01	261136679	144.01	SPOC	S1	5	P	5.65	249	6.2666	3.093	1.9	Huang PI Mensae b Accepted	0	HD 39091 (pi Mensae), RV confirmed planet pi Mensae b (Huang et al. 2018).	94
62483237.01	62483237	139.01	SPOC	S1	1	CPC	10.55	1367	11.0584	2.137	2.7		1	Kris Helminiak observed a full transit in V on 20181031 using deep exposures that saturated the target star. One potential 80 ppt NEB was found at his T5 (73", PA=108, dmag=6.93), but is apparently too faint and shallow to have caused the TESS detection. The next observations should check to see if the potential NEB returns, but otherwise this field has been cleared. Also, attempt to confirm the event on target.	101
382302241.01	382302241	151.01	QLP	S1	5	NEB	11.794	1714	1.5978	1.490	3.6		1	Kevin Collins reduced an LCO full transit observation on 20180914 in Z-band and found no event on target and. However, the dmag = 4.58 star 27" to the NNE (PA=17 degrees) shows a 175 ppt (peak depth) event, which if fully blended in the TESS aperture would produce a 2.5 ppt event. The NEB is probably not fully blended, so the TESS detected depth of 1.7 ppt seems reasonable.	95
403287048.01	403287048	152.01	QLP	S1	5	NEB	11.948	3667	7.0078	2.680	8.1		1	Chris Stockdale observed a full transit on 20181006 in Rc and found a convincing 100 ppt NEB with dmag=3.7 that is 43" WNW of the target. The duration, timing, and depth in the NEB are consistent with the 3.7 ppt event detected by TESS on the brighter target star. The neighbor star is TIC 403287050 at RA=332.332031 (22:09:19.687) and Dec=-62.330091 (-62:19:48.33) and has Tmag=14.7 and Vmag=15.345. Gaia gives Teff=5841+/-400 and Rstar=1.20+/-0.15. Using the lower end of the planet radius, Rp=35.85 Re, so this TOI is an NEB rather than an NPC (nearby planet candidate). Retired as an NEB.	199
220518305.01	220518305	156.01	QLP	S1	5	SB1	10.721	18638	14.1489	3.765	21.8		0	Warm Jupiter candidate with a period of about 14 days. Avi's CHIRON RV data show this target to be an EB. Expired as SB1.	58
231081369.01	231081369	153.01	QLP	S1	5	BEB	12.217	6340	7.6323	1.656			5	Chris Stockdale observed an ingress plus 50-75% on 20180921 in Rc and found no NEBs down to dmag=5.5 and found no clear detection of the predicted 6 ppt transit on target within a scatter of ~4 ppt. Chris Stockdale observed a near full transit on 20181014 in Rc resulting in a 4.2 ppt event on target in an r=6" aperture. An alternate interpretation of the light curve is sinusoidal variations given the pre-ingress OOT variations. No obvious NEBs were found	62
389525208.01	389525208	154.01	QLP	S1	5	NEB	10.493	5000	6.7471	2.343	41.3		1	Phil Evans observed a full transit on 20180915 in Rc and found a mid-transit 80 ppt flux decrement in Rc in a dmag=3.9 (Vmag=14.7) star 21" to the NW (PA=297). Retired as NEB.	44



SG1 TOI List Totals

TFOP SG1 TOTALS		TOTAL	S1	S2	S3	S4	S5	S6	S7	S8	S9	S10	S11	S12	S13
Total SG1 TOIs (excludes KPs and STPCs)		865	54	170	49	59	74	89	89	107	99	39	18	18	0
Level 1		296	12	47	28	27	28	33	22	33	41	12	6	7	0
TOIs with 1+ Observations		445	46	132	33	30	29	32	26	53	37	20	6	1	0
Level 1		141	11	39	16	12	8	8	13	13	12	6	3	0	0
TOIs with 0 Observations (excludes KPs and STPCs)		420	8	38	16	29	45	57	63	54	62	19	12	17	0
Level 1		155	1	8	12	15	20	25	9	20	29	6	3	7	0
SG1 Retired (excludes KPs and STPCs)		241	40	59	17	22	19	16	13	32	16	6	1	0	0
Level 1		72	8	15	11	8	5	6	6	8	4	1	0	0	0
SG1 Active		624	14	111	32	37	55	73	76	75	83	33	17	18	0
Level 1		224	4	32	17	19	23	27	16	25	37	11	6	7	0
Total SG1 Observations		1106	186	349	83	51	61	60	85	89	65	39	12	1	0
Total SG1 Submitted Observations		1081	186	349	83	51	61	60	85	89	65	39	12	1	0
Level 1		366	54	119	42	18	16	16	36	26	24	10	5	0	0
Total Pending Observations		25	0	0	0	0	0	0	0	0	0	0	0	0	0
Planets (excludes KPs)		27	9	12	0	2	1	0	1	1	1	0	0	0	0
P (Confirmed Planet)		22	8	8	0	2	1	0	1	1	1	0	0	0	0
Level 1		9	2	5	0	2	0	0	0	0	0	0	0	0	0
VP (Validated Planet)		5	1	4	0	0	0	0	0	0	0	0	0	0	0
Level 1		4	1	3	0	0	0	0	0	0	0	0	0	0	0
KP (Known Planet)		132	20	24	10	13	1	16	11	4	10	13	10	0	0
Elevated to Other SGs		223	27	72	18	7	12	15	15	23	19	10	5	0	0
VPC+ (SG1 Verified Achromatic PC)		36	10	8	3	1	1	0	4	2	3	2	2	0	0
Level 1		11	2	2	2	0	0	0	1	0	2	1	1	0	0
VPC (SG1 Verified PC on target)		108	11	38	6	2	7	7	5	12	12	6	2	0	0
Level 1		23	2	7	2	1	1	0	2	3	3	1	1	0	0
VPC- (SG1 Verified PC w/contaminated aperture)		5	0	1	0	0	0	0	1	1	1	0	0	0	0
Level 1		1	0	1	0	0	0	0	0	0	0	0	0	0	0
CPC (SG1 Cleared PC - no NEBs)		39	3	14	7	1	2	3	4	2	2	1	0	0	0
Level 1		32	3	10	5	1	1	3	4	2	2	1	0	0	0
CPC- (SG1 Cleared PC outside target PSF)		4	0	0	2	0	0	1	0	1	0	0	0	0	0
Level 1		3	0	0	2	0	0	1	0	0	0	0	0	0	0
PPC (SG1 Promising PC)		31	3	11	0	3	2	4	1	5	1	0	1	0	0
Level 1		15	1	5	0	3	1	1	1	1	1	0	1	0	0
Not Elevated to Other SGs		472	4	63	22	31	46	62	66	54	69	24	13	18	0
PC (still a Planet Candidate)		453	2	55	22	31	46	59	65	52	67	24	12	18	0
Level 1		165	0	14	13	14	20	23	11	20	31	9	3	7	0
PNEB (Possible NEB)		1	0	0	0	0	0	0	0	1	0	0	0	0	0
Level 1		0	0	0	0	0	0	0	0	0	0	0	0	0	0
STPC (Single Transit PC)		18	2	8	0	0	0	3	1	1	2	0	1	0	0
Level 1		3	0	2	0	0	0	1	0	0	0	0	0	0	0
False Positives		160	16	31	8	19	15	15	8	30	12	5	1	0	0
NEB (Nearby EB)		63	3	9	1	13	6	9	3	15	4	0	0	0	0
Level 1		22	1	0	1	6	3	3	3	4	1	0	0	0	0
NPC (Nearby PC)		12	3	3	1	1	1	1	0	1	1	0	0	0	0
Level 1		1	0	0	0	0	1	0	0	0	0	0	0	0	0
BEB (chromaticity)		21	2	8	3	1	1	1	2	1	1	0	1	0	0
Level 1		0	0	0	0	0	0	0	0	0	0	0	0	0	0
BD (confirmed Brown Dwarf)		2	0	0	0	0	0	0	1	1	0	0	0	0	0
Level 1		0	0	0	0	0	0	0	0	0	0	0	0	0	0
EB (too deep)		14	0	7	3	2	0	1	1	0	0	0	0	0	0
Level 1		1	0	0	0	0	0	1	0	0	0	0	0	0	0
SB1 (too large RV)		32	5	2	0	2	6	1	1	6	5	4	0	0	0
Level 1		0	0	0	0	0	0	0	0	0	0	0	0	0	0
SB2 (multi-lined)		9	2	2	0	0	0	1	0	3	0	1	0	0	0
Level 1		0	0	0	0	0	0	0	0	0	0	0	0	0	0
FA (false alarm)		7	1	0	0	0	1	1	0	3	1	0	0	0	0
Level 1		5	0	0	0	0	1	1	0	3	0	0	0	0	0

- ◆ Chart linked on TESS Wiki
- ◆ 865 (L1=296) SG1 TOIs (excluding KPs and STPCs)
- ◆ 1100+ (L1=375+) SG1 submitted observations
 - ~60% LCO observers
 - ~40% from non-LCO observers
- ◆ ~50% of TOIs have 1+ observations
- ◆ ~20% FPs
- ◆ See SG1 Related Posters:
 - Poster 5 – McLeod et al.
 - Poster 7 – Narita et al.
 - Poster 9 – Waalkes et al.
 - Poster 11 – Fukui et al.
 - Poster 12 – Kielkopf et al.
 - Poster 19 – Palles et al.
 - Poster 54 – Kafka et al.

- ◆ Interested in a specific TOI?
 - *Let me know early on and I can boost SG1 priority*
- ◆ Recommend including constraints from follow-up in paper
 - *Size of follow-up aperture (in arcsec) used to confirm event on target*
 - *FWHM of stars in follow-up imagery*
 - *Include light curves in global model*
 - *Reduce ephemeris uncertainty by up to ~90%*
 - *Verify TESS deblending factor via follow-up depths*
 - *Verify achromatic transit via multi-band follow-up*
 - *TTV constraints*
- ◆ TOIs with depths less than ~1.0 ppt
 - *The CPC status rules out NEBs*
- ◆ We can help write SG1 observations and data reduction section and provide interpretation
- ◆ TFOP observers that contributed useful data, including ruling out of a false positive, should be offered coauthorship, even if you later confirmed the planet mass with RVs.
 - *Check ExoFOP-TESS and SG1 Google Sheet*
 - *Ask me to help identify appropriate SG1 coauthors*

